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COVER ILLUSTRATION:

Instability domains in the upper part of the Hertzsprung-Russel Diagram. Slowly Pulsating B stars are unstable only to nonradial high-order gravity modes. A similar instability was found for massive stars in addition to the β Cep-type instability. For the classical instability strip, only the Blue Edge is shown. A few evolutionary tracks for the indicated values of M/M_{\odot} are shown. The location of γ Dor variables and of stars with solar-like oscillation is also shown. The position of the present Sun is marked by the symbol \odot . For more information please see the article by A. A. Pamyatnykh on p. 10.

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